

Energy staggering parameters in nuclear magnetic rotational bands*

Wu-Ji Sun(孙无忌)¹⁾ Jian Li(李剑)¹⁾

College of Physics, Jilin University, Changchun 130012, China

Abstract: This study presents the systematics of energy staggering for magnetic rotational bands with $M1$ and $E2$ transition properties, which are strictly consistent with the features of good candidates of magnetic rotational bands in the $A \sim 80, 110, 130$, and 190 mass regions. The regularities exhibited by these bands with respect to the staggering parameter, which increases with increasing spin, are in agreement with the semiclassical description of shears mechanism. Moreover, the abnormal behaviour in the backbend regions or close to band termination has also been discussed. Taking the magnetic dipole bands with same configuration in three $N = 58$ isotones, i.e., ^{103}Rh , ^{105}Ag , and ^{107}In , as examples, the transition from chiral to magnetic rotation with the proton number approaching $Z = 50$ is presented. Moreover, the self-consistent tilted axis and principle axis cranking relativistic mean-field theories are applied to investigate the rotational mechanism in the dipole band of ^{105}Ag .

Keywords: energy staggering, magnetic rotational band, shears mechanism, cranking relativistic mean-field theory

PACS: 21.60.Jz, 21.30.Fe, 27.60.+j **DOI:** 10.1088/1674-1137/43/9/094104

1 Introduction

Similar to rotational bands observed in molecules, many nuclei have an energy spectrum with a pronounced rotational character, and the study of nuclear rotation has been at the forefront of nuclear structure for several decades. In particular, magnetic rotation (MR) [1], which is an exotic rotational phenomenon observed in weakly deformed or near-spherical nuclei and differs from conventional collective rotation in well-deformed nuclei, has been of great interest since the observation of cascades of magnetic dipole ($M1$) transitions in the region of neutron-deficient Pb isotopes in the 1990s [2–4].

The explanation of MR was given in terms of the shears mechanism [5]. In this interpretation, the magnetic dipole vector in the magnetic rotational bands arising from proton particles (holes) and neutron holes (particles) in high- j orbitals rotates around the total angular-momentum vector. Meanwhile, with increasing spin, the proton particles (holes) and neutron holes (particles) in the high- j orbitals align along the total angular momentum, which reduces the perpendicular component of the magnetic dipole moment. Consequently, a typical property of these bands is the decreasing of the $B(M1)$ values with increasing spin. Experimental indicators for magnetic rotational bands can be summarized as follows [6–10]: 1) a

$\Delta I = 1$ sequence of strong magnetic dipole ($M1$) transitions, corresponding to a reduced transition probability $B(M1) \sim$ a few μ_N^2 , which decrease with increasing spin; 2) weak or absent quadrupole transitions, corresponding to a deformation parameter $|\beta| \lesssim 0.15$, which combined with strong $M1$ transitions results in large $B(M1)/B(E2)$ ratios, $\gtrsim 20 \mu_N^2/(eb)^2$; 3) a smooth variation in the γ -transition energy with angular momentum; 4) a substantial moment of inertia, corresponding to the large ratio of the $\mathcal{J}^{(2)}/B(E2) \gtrsim 100 \text{ MeV}^{-1}(eb)^{-2}$, compared with the values in well-deformed [$\sim 10 \text{ MeV}^{-1}(eb)^{-2}$] or superdeformed [$\sim 5 \text{ MeV}^{-1}(eb)^{-2}$] rotational bands.

The first clear evidence of magnetic rotational bands was provided by the lifetime measurements for four $M1$ -bands in $^{198,199}\text{Pb}$ [11]. Afterwards, an increasing number of magnetic rotational bands were observed not only in the mass region of $A \sim 190$, but also in the $A \sim 60, 80, 110$, and 130 regions. To date, more than 200 magnetic dipole bands spread over 110 nuclides have been observed, which are summarized in the nuclear chart in Fig. 1 based on the review on the observed MR bands [12] and recent observations [13–50]. The blue squares in Fig. 1 represent nuclides with good candidate MR bands, in which $M1$ and $E2$ transition properties are strictly consistent with the features of MR, i.e., decreasing $B(M1)$ values with increasing spin and $B(M1)/B(E2) \gtrsim 20 \mu_N^2/(eb)^2$,

Received 15 April 2019, Revised 26 June 2019, Published online 23 July 2019

* Supported by National Natural Science Foundation of China (11675063, 11205068, 11847310, 11775098)

1) E-mail: jianli@jlu.edu.cn

©2019 Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd

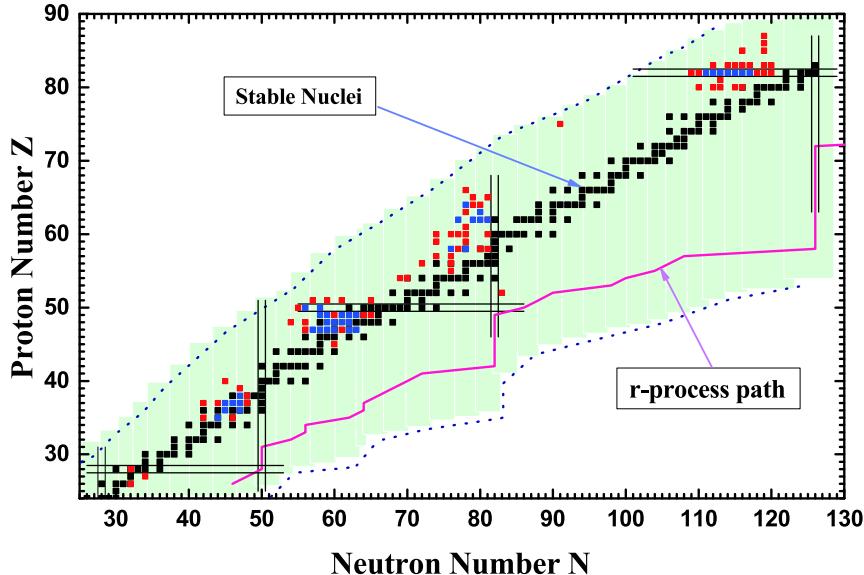


Fig. 1. (color online) Candidate nuclides with magnetic rotation observed in the nuclear chart. Blue squares represent nuclides with good candidate MR bands, in which $M1$ and $E2$ transition properties are strictly consistent with the features of magnetic rotation. Red squares represent other candidates. The corresponding data are taken from Refs. [12–50].

including 50 bands spread over 39 nuclides. The red squares represent other MR candidate nuclides without lifetime measurements.

Signature is the quantum number specifically appearing in a deformed intrinsic system and associated with symmetry under a rotation of 180° about the axis of nuclear rotation. A rotational band with a $\Delta I = 1$ sequence could be divided into two branches classified by the signature quantum number [51–53]. Experimentally, energy staggering between alternate spin states, which is best visualized by the experimental quantity $S(I) = [E(I) - E(I-1)]/2I$, could be observed in many rotational bands and is usually referred to as signature splitting. Theoretically, no signature splitting should be observed in ideal magnetic rotational bands due to the pure individual motion of nucleons (shears mechanism), and the tilted angle generally remains far from 90° in the tilted rotational mode of nuclei [5, 7]. However, it is noted that the experimental energy staggering could be found for some MR candidates, as mentioned in Ref. [54].

In magnetic rotational bands, their energy spectra show rotational-like features with strong $M1$ transitions, and it is noted in Ref. [6] that the rotational-like energy spectra (away from band crossings) approximately follow the pattern of $E(I) - E_0 \propto (I - I_0)^2$, such as in Pb isotopes [2–4], where $E(I) - E_0$ is the relative energy of a state with spin I at energy $E(I)$ to the energy of the band-head state, E_0 . The rotational-like behavior observed could be explained by a semiclassical analysis of the shears mechanism from a residual proton-neutron interaction [55]. Therefore, in the MR bands, $S(I) \propto 1 - (2I_0 + 1)/2I$ are supposed to increase with increasing spin, and signa-

ture splitting, i.e., energy staggering, is not expected. In comparison, for conventional rotation in well-deformed nuclei, their energy spectra follow the pattern $E(I) - E_0 \propto I(I+1)$ [51] based on the simple assumption of a constant moment of inertia, while $S(I)$ values are constant at various spins.

Thus, it is interesting to systematically study energy staggering $S(I)$ of magnetic rotational bands, especially the ones with suitable electromagnetic transition properties. The present work investigates systematic behaviors of $S(I)$ in good candidates of MR bands. Furthermore, taking the dipole bands with same configuration in ^{103}Rh , ^{105}Ag , and ^{107}In as examples, the characteristics of $S(I)$ in different rotational modes would be compared.

2 Energy staggering systematics of magnetic rotational bands

In Fig. 2, the experimental energy staggering $S(I)$ for the good candidate MR bands in the mass $A \sim 80$, 110, 130, and 190 regions are shown, which include 50 bands spread out in 39 nuclides marked with blue squares in Fig. 1. In the 110, 130, and 190 mass regions, the MR bands in even-even, odd-odd, and odd- A nuclei are presented. The corresponding data are taken from the previous review [12] and recent observations [15–29, 56–58].

Figure 2 shows that for almost all the MR candidates, the experimental energy staggering $S(I)$ tend to increase with increasing spin except for the backbend regions (or band crossing), in agreement with the semi-classical for-

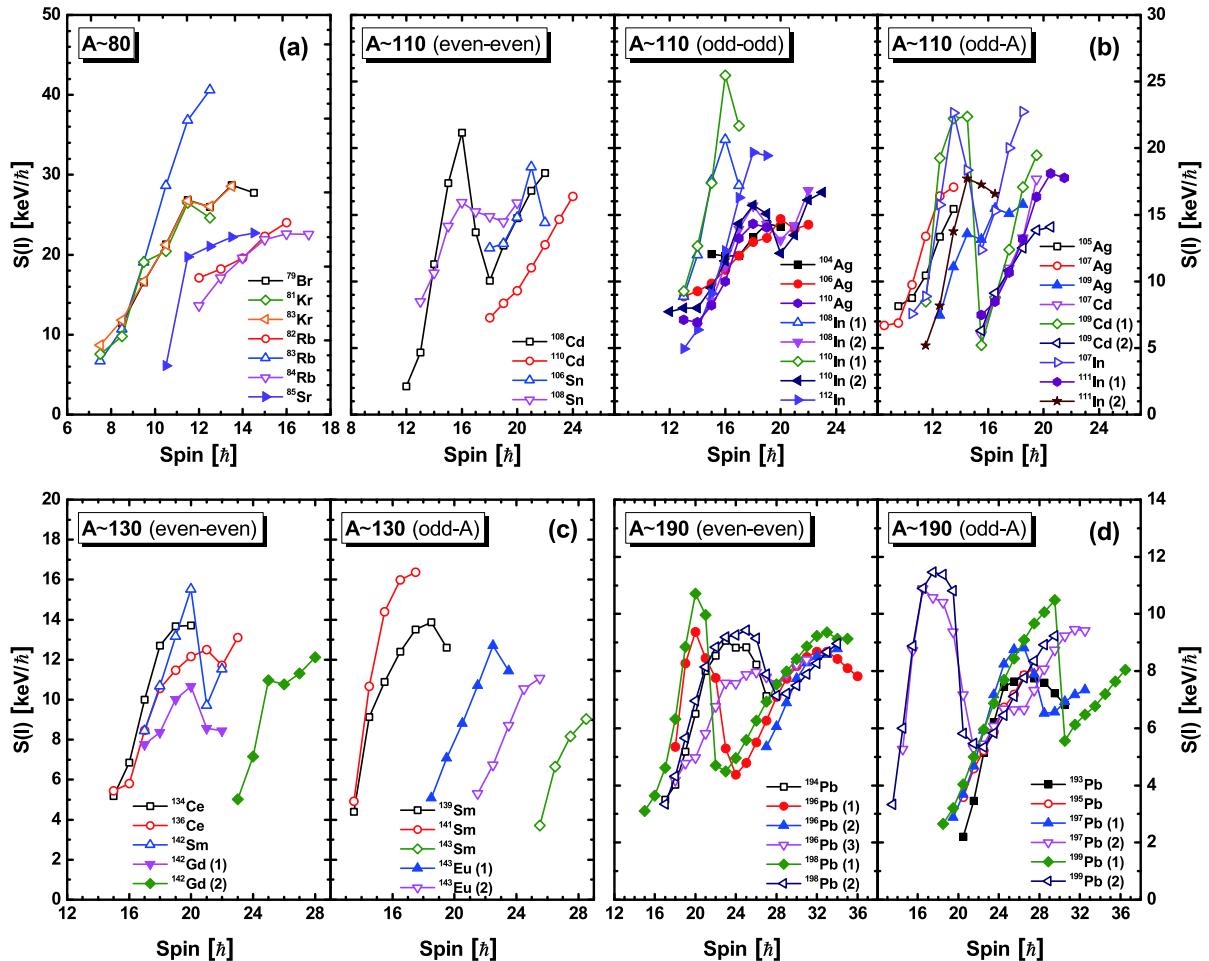


Fig. 2. (color online) Experimental energy staggering $S(I)$ as a function of spin I for good candidate MR bands in the mass number $A \sim 80$ (a), 110 (b), 130 (c) and 190 (d) regions. In the 110, 130, and 190 mass regions, MR bands in even-even, odd-odd, and odd- A nuclei are presented separately. The filled and open symbols depict positive parity bands and negative parity bands, respectively. The corresponding data are taken from the previous review Ref. [12] and recent observations [15–29, 56–58].

mula for the rotational-like properties of MR, i.e., the deduced term $S(I) \propto 1 - (2I_0 + 1)/2I$, and differently from the semi-classical formula for conventional rotation in well-deformed nuclei. For example, for the MR bands of ^{79}Br , $^{81,83}\text{Kr}$, $^{82,83,84}\text{Rb}$, and ^{85}Sr in the 80 mass region, $S(I)$ clearly exhibits an increasing pattern as the spin increases. In Ref. [59], the semiclassical approach using two blades (particle and hole) interacting with an effective force is proposed, and this simple scenario can account for the important features of the shears bands in neutron-deficient Pb nuclei and other mass regions. In fact, the moment of inertia in the shears mechanism would decrease with the closing of the shears blades [7, 59]. The behavior where the kinematic moment of inertia $\mathcal{J}^{(1)} (= 1/[2S(I)])$ decreases with angular momentum is clearly demonstrated in Ref. [59], which further proves the increasing feature of $S(I)$ with spin. However, $S(I)$ for the MR band in ^{108}Sn [60] shown in Fig. 2(b) is an exception, i.e., in the higher spin region of the band, $S(I)$ be-

gins to decrease, while $B(M1)$ still shows a decreasing tendency.

The $S(I)$ values in the backbend region show a decreasing trend with increasing spin for the MR bands in the 110, 130, and 190 mass regions. The backbend phenomenon is usually interpreted as the decoupling of a pair of particles from the rotating nuclear core and the subsequent rotational alignment of their angular momenta along the rotation axis [61]. The backbend would result in an increasing moment of inertia caused by the shears opening due to the gradual alignment of the particles, and therefore $S(I)$ shows the decreasing trend in the backbend regions of MR bands. For example, the $S(I)$ of the negative-parity MR band of ^{199}Pb decreases from 11.5 to 5.3 keV/hbar in the backbend region [62], as shown in Fig. 2. Here, the configurations of the negative-parity MR band in ^{199}Pb are $\pi(h_{9/2}i_{13/2})_{K=11^-} \otimes vi_{13/2}^{-1}$ before the backbend ($I \leq 17.5\hbar$) and $\pi(h_{9/2}i_{13/2})_{K=11^-} \otimes vi_{13/2}^{-3}$ after the backbend ($I \geq 23.5\hbar$), respectively [62]. After the rota-

tional alignment of two $i_{13/2}$ neutron holes, the shears open up to 90° coupling, and a new shears band starts building up.

In addition, it is easy to see in Fig. 2 that the $S(I)$ values in higher mass regions are relatively smaller compared to those in lower mass regions. In the $A \sim 80$ region, $S(I)$ varies from 8 to 42 keV/ \hbar and it varies from 2 to 25 keV/ \hbar in the $A \sim 110$ region. Then, in the $A \sim 130$ region, $S(I)$ varies from 4 to 16 keV/ \hbar , and finally it varies from 2 to 12 keV/ \hbar in the $A \sim 190$ region. This behavior is closely related to the mass dependence of the moment of inertia for these bands, as discussed in Ref. [55]. Moreover, $S(I)$ is also related to the corresponding valence particle-hole configuration. For example, for the positive-parity band in $^{108,110,112}\text{In}$ [18, 19] with the same configuration $\pi g_{9/2}^{-1} \otimes \nu h_{11/2}^2(g_{7/2}/d_{5/2})^1$, the $S(I)$ values of these bands are quite similar, as shown in Fig. 2(b). Meanwhile for the two bands with different configurations in the same nuclide ^{108}In , there is a clear discrepancy in the $S(I)$ values.

However, some MR bands in Fig. 2 show abnormal behaviours of $S(I)$ for the band termination (high spin) region. For example, a sudden decline in $S(I)$ occurs at the highest spin ($I = 23.5\hbar$) of the MR band of ^{143}Eu in Fig. 2(c), or $S(I)$ decreases in the high-spin region ($I > 26.5\hbar$) of ^{193}Pb in Fig. 2(d). Similar behaviors are observed in ^{104}Ag , $^{110,111,112}\text{In}$, ^{106}Sn , ^{136}Ce , ^{139}Sm , ^{142}Gd , and $^{194,196,198}\text{Pb}$. The decreasing tendency of $S(I)$ in the high-spin region of the bands, in some cases a sudden decline at the highest spin, are mostly caused by the abrupt change of the configuration, which is similar to the pattern of $S(I)$ in the backbend region. For the MR band of ^{109}Ag in Fig. 2(b), $S(I)$ shows a staggering pattern in the high-spin region ($I \geq 15.5\hbar$), which is also found in ^{79}Br , $^{81,83}\text{Kr}$, $^{105,106,107}\text{Ag}$, and ^{136}Ce , as shown in Fig. 2. In Ref. [63], it is noted that the competition and interaction with collective rotations happen in the high-spin region of the bands with a certain configuration, as aligned states

do not have the maximum spin.

In all the presented MR bands, the increasing tendency of energy staggering $S(I)$ with increasing spin before and after the backbend are observed. The systematic study of the energy staggering parameter in MR bands shows a common behavior of $S(I)$, implying that $S(I)$ could be a potential indicator for MR. Therefore, it is necessary to test and compare with other exotic rotations, such as chiral rotation [64]. Taking the $M1$ bands in ^{103}Rh [65], ^{105}Ag [20, 66, 67], and ^{107}In [24, 68] with the same configuration $\pi g_{9/2}^{-1} \otimes \nu h_{11/2}^1(g_{7/2}/d_{5/2})^1$ as examples, the extracted experimental $S(I)$ values are shown and compared in Fig. 3.

3 Magnetic rotation in the $A \sim 110$ region

3.1 The transition from chiral rotation to magnetic rotation

The two $M1$ bands of ^{103}Rh have been proposed as chiral doublet bands [65]. Multiple chiral doublet bands with the same configuration $\pi g_{9/2}^{-1} \otimes \nu h_{11/2}^1(g_{7/2}/d_{5/2})^1$ were observed in ^{103}Rh , and only the "yrast" chiral doublet bands are presented here. The $S(I)$ values of those two chiral partner bands in Fig. 3(a) remain almost constant with a little staggering, and $B(M1)/B(E2)$ ratios show typical values and behaviors of chiral doublet bands, as suggested in Ref. [65].

For the $M1$ band of ^{107}In , the $S(I)$ values increase with increasing spin before and after the backbend. The behaviors of the experimental data in this band, including $B(M1)/B(E2)$ ratios and $B(M1)$ values, are consistent with ideal MR bands, as suggested in Ref. [24].

For the $M1$ band of ^{105}Ag , the $S(I)$ values of the $M1$ band tend to increase with increasing spin in the lower spin region, which is consistent with the behavior in ideal MR bands. In the higher spin region ($I > 15.5\hbar$), $S(I)$ shows noticeable staggering. The staggering of $S(I)$ is usually considered as a sign of the collective rotation,

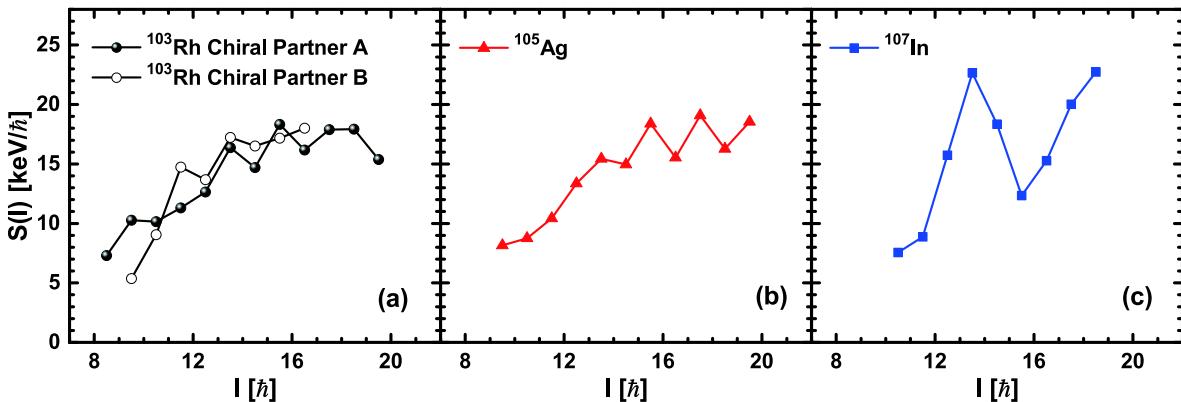


Fig. 3. (color online) Experimental $S(I)$ as a function of spin I for the negative parity $M1$ bands of ^{103}Rh (a), ^{105}Ag (b), and ^{107}In (c) with the same configuration $\pi g_{9/2}^{-1} \otimes \nu h_{11/2}^1(g_{7/2}/d_{5/2})^1$. The corresponding data are taken from Refs. [20, 24, 65–68].

which indicates a competition between the shears mechanism and collective motion.

The single-particle Routhians for the proton at the top of the $\pi g_{7/2}$ shell is close to the $\pi g_{9/2}$ proton shell. Considering that the contribution from collective motion is increasing, there can be a mixing of the configuration $\pi g_{9/2}^{-1} \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$ and the configuration $\pi g_{7/2}^1 \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$ at high-spin states.

Based on the above statements, the shears mechanism does not seem to dominate in the $M1$ bands of ^{103}Rh and ^{105}Ag , and there is an obvious transition from chiral rotation to magnetic rotation in the $A \sim 110$ region, when the proton number is approaching $Z = 50$. Moreover, the quadrupole deformation in previous TAC calculations are 0.26, 0.19, and 0.12 for ^{103}Rh [65], ^{105}Ag [69] and ^{107}In [24] with the same configuration $\pi g_{9/2}^{-1} \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$, respectively. This probably indicates that the deformation of nuclei becomes smaller with increasing Z number towards $Z = 50$. Similarly to Refs. [21, 70], the increasing importance of the shears mechanism and decreasing contribution of collective rotation with increasing proton number towards $Z = 50$ in the $A \sim 110$ region have also been presented and discussed. Thus, the systematic study of the different rotational modes in the $A \sim 110$ region is an interesting question.

3.2 The magnetic dipole band in ^{105}Ag

During the past few decades, the relativistic mean-field (RMF) theory has been a great success in describing properties of nuclei and many nuclear phenomena [71–73]. The principle axis cranking relativistic mean-field (PAC-RMF) theory has been used to describe collective rotational motion in deformed nuclei [74]. Based on the RMF theory, the tilted axis cranking relativistic mean-field (TAC-RMF) theory has been developed for describing nuclear magnetic and antimagnetic rotational modes [9, 10]. The cranking RMF model with arbitrary orientation of the rotational axis, i.e., three-dimensional cranking, has been developed and applied for the magnetic rotation in ^{84}Rb [75]. Recently, the three-dimensional TAC-RMF theory with the point-coupling interaction has been used to investigate multiple chirality in nuclear rotation [76, 77]. The two-dimensional cranking RMF theory based on the meson exchange [78] and the point-coupling interactions [79, 80] has also been established and applied successfully to describe magnetic rotation in $A \sim 60, 80, 130$, and 190 regions [9, 10], and especially the 110 region [28, 80–86].

To further examine the rotational mechanism, as well as the staggering pattern of $S(I)$ for the band in ^{105}Ag , the TAC-RMF calculations for the lower spin region of the dipole band in ^{105}Ag with the configuration $\pi g_{9/2}^{-1} \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$ and the principle axis cranking RMF

(PAC-RMF) calculations for the higher spin region of the band with the configuration $\pi g_{7/2}^1 \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$ in different signatures have been performed. The point-coupling interaction PC-PK1 [87] was used, and pairing correlations were neglected. The Dirac equation for the nucleons is solved in a three-dimensional harmonic oscillator basis, and a basis of ten major oscillator shells is adopted. The calculated rotational excitation energies, total angular momenta and $B(M1)$ values in comparison with the corresponding data [20, 66, 67] are shown in Fig. 4. The results of TAC-RMF calculations for the lower spin region and PAC-RMF calculations for the higher spin region in the $M1$ band are displayed. The experimental rotational frequency can be extracted as in Ref. [88]: $\hbar\omega_{\text{exp}} = \frac{1}{2}[E_\gamma(I+1 \rightarrow I) + E_\gamma(I \rightarrow I-1)]$.

Figure 4(a) shows that the calculated excitation ener-

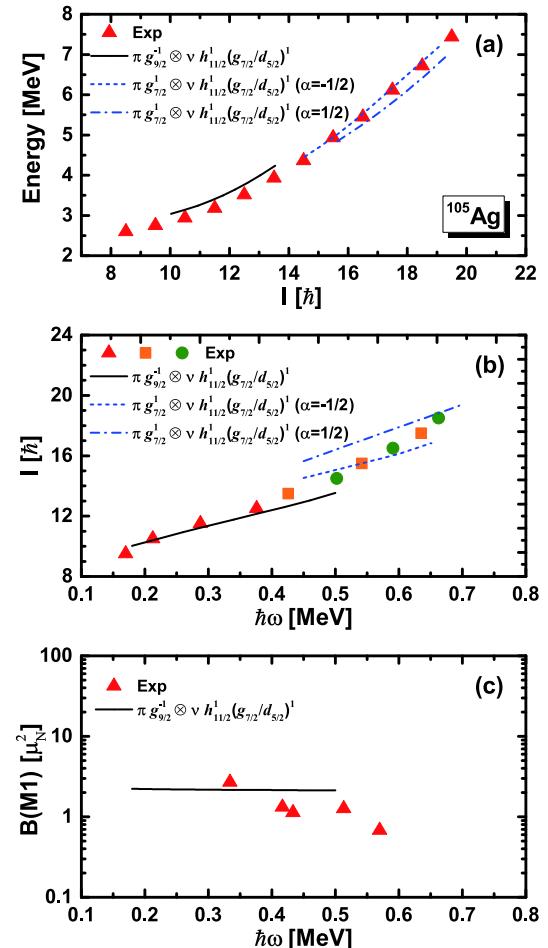


Fig. 4. (color online) Rotational excitation energies (a) as a function of the total angular momentum, total angular momenta (b), and $B(M1)$ values (c) as functions of the rotational frequency, for the configurations $\pi g_{9/2}^{-1} \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$ in TAC-RMF calculations and $\pi g_{7/2}^1 \otimes v h_{11/2}^1 (g_{7/2}/d_{5/2})^1$ in PAC-RMF calculations. The corresponding data are taken from Refs. [20, 66, 67].

gies of the band in both TAC-RMF and PAC-RMF calculations are both in good agreement with experimental data. In Fig. 4(b), for the lower spin region, the TAC-RMF results reproduce the data well, which further supports the significant contribution from shears mechanism. For the higher spin region, the PAC-RMF results are in reasonable agreement with the data, indicating that the increasing contribution from the collective motion results from the intruder configuration $\pi g_{7/2}^1 \otimes vh_{11/2}^1(g_{7/2}/d_{5/2})^1$. Moreover, the experimental $I - \omega$ plot is sandwiched between the corresponding plots for the configuration $\pi g_{9/2}^{-1} \otimes vh_{11/2}^1(g_{7/2}/d_{5/2})^1$ and $\pi g_{7/2}^1 \otimes vh_{11/2}^1(g_{7/2}/d_{5/2})^1$, further supporting the mixing of these two configurations at high-spin states.

In Fig. 4(c), the calculated $B(M1)$ values are shown to be in agreement with the data, however they do not decrease much with increasing rotational frequency, corresponding to a small decline in the shears angle. As discussed in Ref. [78], the tilted angle θ of the orientation of the angular velocity with respect to the principal axis of the density distribution is determined self-consistently in the TAC-RMF calculations. With the rotational frequency increasing from 0.18 to 0.50 MeV/ \hbar , the tilted angle of the proton angular momentum θ_π changes from 5° to 10°, and the tilted angle of neutron angular momentum θ_ν decreases from 81° to 71°. The shears angle between these two blades decreases from 76° to 61°, i.e., the proton and neutron angular momenta align toward each other with increasing rotational frequency, which exhibits an apparent shears mechanism. The shears angle only decreases by a small amount due to the relatively high contribution from collective motion. In general, the

TAC-RMF calculations support the MR interpretation for the lower spin region of the band.

4 Conclusions

In summary, 50 bands spread over 39 nuclides with $M1$ and $E2$ transition properties, which are strictly consistent with the features of MR, have been selected, and a systematic study of energy staggering parameter $S(I)$ in these bands has been performed. The present study shows that $S(I)$ values increase with increasing spin for all the bands before and after backbends, which could be explained by the simple semiclassical description of the shears mechanism. This could be treated as an indicator for MR, and also requires more investigations. Moreover, the behaviors of $S(I)$ in the backbend regions or close to band termination have been discussed. In addition, the $M1$ bands in three $N = 58$ isotones, i.e., ^{103}Rh , ^{105}Ag , and ^{107}In , with the same configuration $\pi g_{9/2}^{-1} \otimes vh_{11/2}^1(g_{7/2}/d_{5/2})^1$ are taken as examples to further examine the staggering behaviours of $S(I)$ in different rotational modes. It is suggested that there is a transition from chiral rotation to magnetic rotation with the proton number approaching $Z = 50$, due to the competition between the collective motion and the shears mechanism. Furthermore, the TAC-RMF and PAC-RMF calculations have been performed, and the rotational modes in ^{105}Ag are clearly shown.

The authors would like to thank Dr. P. W. Zhao and Y. K. Wang for helpful discussions and collaboration during the completion of this work.

References

- 1 S. Frauendorf, J. Meng, and J. Reif, in *Proceedings of the Conference on Physics From Large-Ray Detector Arrays*, Vol. II of Report LBL35687, edited by M. A. Deleplanque (University of California, Berkeley, 1994), p. 52
- 2 R. M. Clark, R. Wadsworth, E. S. Paul et al, *Phys. Lett. B*, **275**: 247 (1992)
- 3 A. Kuhnert, M. A. Stoyer, J. A. Becker et al, *Phys. Rev. C*, **46**: 133 (1992)
- 4 G. Baldsiefen, H. Hübel, D. Mehta et al, *Phys. Lett. B*, **275**: 252 (1992)
- 5 S. Frauendorf, *Nucl. Phys. A*, **557**: 259c (1993)
- 6 R. M. Clark and A. O. Macchiavelli, *Ann. Rev. Nucl. Part. Sci.*, **50**: 1 (2000)
- 7 S. Frauendorf, *Rev. Mod. Phys.*, **73**: 463 (2001)
- 8 H. Hübel, *Prog. Part. Nucl. Phys.*, **54**: 1 (2005)
- 9 J. Meng, J. Peng, S. Q. Zhang et al, *Front. Phys.*, **8**: 55 (2013)
- 10 J. Meng, S. Q. Zhang, and P. W. Zhao, in *Relativistic Density Functional for Nuclear Structure*, Vol. 10, edited by J. Meng, (World Scientific, Singapore, 2016)
- 11 R. M. Clark, S. J. Asztalos, G. Baldsiefen et al, *Phys. Rev. Lett.*, **78**: 1868 (1997)
- 12 Amita, A. K. Jain, and B. Singh, *At. Data Nucl. Data Tables*, **74**: 283 (2000)
- 13 J. Li, C. Y. He, Y. Zheng et al, *Phys. Rev. C*, **88**: 014317 (2013)
- 14 Y. Zheng, J. Li, J. J. Liu et al, *J. Phys. G*, **42**: 085108 (2015)
- 15 N. Kumar, S. Kumar, S. K. Mandal et al, *Eur. Phys. J. A*, **53**: 25 (2017)
- 16 B. Das, P. Datta, S. Chatopadhyay et al, *Phys. Rev. C*, **98**: 014326 (2018)
- 17 S. Rajbanshi, R. Raut, H. Pai et al, *Phys. Rev. C*, **98**: 061304(R) (2018)
- 18 C. J. Chiara, D. B. Fossan, V. P. Janzen et al, *Phys. Rev. C*, **64**: 054314 (2001)
- 19 T. Trivedi, R. Palit, J. Sethi et al, *Phys. Rev. C*, **85**: 014327 (2012)
- 20 A. Y. Deo, S. B. Patel, S. K. Tandel et al, *Phys. Rev. C*, **73**: 034313 (2006)
- 21 S. H. Yao, H. L. Ma, L. H. Zhu et al, *Phys. Rev. C*, **89**: 014327 (2014)
- 22 P. Datta, S. Roy, S. Pal et al, *Phys. Rev. C*, **78**: 021306 (2008)
- 23 D. Choudhury, R. Palit, P. Singh et al, *Phys. Rev. C*, **91**: 014318 (2015)
- 24 D. Negi, T. Trivedi, A. Dhal et al, *Phys. Rev. C*, **81**: 054322 (2010)
- 25 C. Y. He, X. Q. Li, L. H. Zhu et al, *Phys. Rev. C*, **83**: 024309 (2011)

- 26 S. Rajbanshi, S. Ali, A. Bisoi et al, *Phys. Rev. C*, **94**: 044318 (2016)
- 27 S. Rajbanshi, A. Bisoi, S. Nag et al, *Phys. Rev. C*, **89**: 014315 (2014)
- 28 C. B. Li, J. Li, X. G. Wu et al, *Nucl. Phys. A*, **892**: 34 (2012)
- 29 S. Rajbanshi, A. Bisoi, S. Nag et al, *Phys. Rev. C*, **90**: 024318 (2014)
- 30 D. Negi, T. Trivedi, A. Dhal et al, *Phys. Rev. C*, **85**: 057301 (2012)
- 31 C. M. Petrache, Q. B. Chen, S. Guo et al, *Phys. Rev. C*, **94**: 064309 (2016)
- 32 T. Bhattacharjee, S. Chanda, S. Bhattacharyya et al, *Nucl. Phys. A*, **825**: 16 (2009)
- 33 S. Kaim, C. M. Petrache, A. Gargano et al, *Phys. Rev. C*, **91**: 024318 (2015)
- 34 N. T. Zhang, Y. H. Zhang, X. H. Zhou et al, *Phys. Rev. C*, **84**: 057302 (2011)
- 35 S. Y. Wang, D. P. Sun, B. Qi et al, *Phys. Rev. C*, **86**: 064302 (2012)
- 36 A. Y. Deo, R. Palit, Z. Naik et al, *Phys. Rev. C*, **79**: 067304 (2009)
- 37 R. Garg, S. Kumar, M. Saxena et al, *Phys. Rev. C*, **92**: 054325 (2015)
- 38 Y. Y. Cheng, S. Q. Zhang, X. Q. Li et al, *Phys. Rev. C*, **89**: 054309 (2014)
- 39 M. G. Procter, D. M. Cullen, C. Scholey et al, *Phys. Rev. C*, **81**: 054320 (2010)
- 40 K. Auranen, J. Uusitalo, S. Juutinen et al, *Phys. Rev. C*, **91**: 024324 (2015)
- 41 D. J. Hartley, E. P. Seyfried, W. Reviol et al, *Phys. Rev. C*, **78**: 054319 (2008)
- 42 H. J. Li, B. Cederwall, M. Doncel et al, *Phys. Rev. C*, **93**: 034309 (2016)
- 43 A. D. Ayangeakaa, S. Zhu, R. V. F. Janssens et al, *Phys. Rev. C*, **91**: 044327 (2015)
- 44 H. Pai, G. Mukherjee, S. Bhattacharyya et al, *Phys. Rev. C*, **85**: 064313 (2012)
- 45 T. Zerrouki, C. M. Petrache, R. Leguillon et al, *Eur. Phys. J. A*, **51**: 50 (2015)
- 46 A. Herzān, S. Juutinen, K. Auranen et al, *Phys. Rev. C*, **96**: 014301 (2017)
- 47 D. Steppenbeck, R. V. F. Janssens, S. J. Freeman et al, *Phys. Rev. C*, **85**: 044316 (2012)
- 48 D. A. Torres, F. Cristancho, L.-L. Andersson et al, *Phys. Rev. C*, **78**: 054318 (2008)
- 49 D. Yuan, Y. Zheng, D. Zhou et al, *Hyperfine Interact.*, **180**: 49 (2008)
- 50 P. Agarwal, S. Kumar, S. Singh et al, *Phys. Rev. C*, **76**: 024321 (2007)
- 51 A. Bohr and B. R. Mottelson, *Nuclear Structure Volume II: Nuclear Deformation* (W. A. Benjamin, Inc., 1975)
- 52 Y. S. Chen, S. Frauendorf, and G. A. Leander, *Phys. Rev. C*, **28**: 2437 (1983)
- 53 Victor Velázquez, Jorge G. Hirsch, and Yang Sun, *Nucl. Phys. A*, **686**: 129 (2001)
- 54 Amita, A. K. Jain, A. Goel, and B. Singh, *Pramana*, **53**: 463 (1999)
- 55 A. O. Macchiaielli, R. M. Clark, P. Fallon et al, *Phys. Rev. C*, **57**: R1073 (1998)
- 56 P. Banerjee, S. Ganguly, M. K. Pradhan et al, *Phys. Rev. C*, **83**: 024316 (2011)
- 57 A. A. Pasternak, E. O. Lieder, R. M. Lieder et al, *Eur. Phys. J. A*, **37**: 279 (2008)
- 58 B. Das, N. Rather, P. Datta et al, *Phys. Rev. C*, **95**: 051301 (2017)
- 59 A. O. Macchiaielli, R. M. Clark, M. A. Deleplanque et al, *Phys. Rev. C*, **58**: 3746 (1998)
- 60 D. G. Jenkins, R. Wadsworth, J. A. Cameron et al, *Phys. Rev. Lett.*, **83**: 500 (1999)
- 61 F. S. Stephens and R. S. Simon, *Nucl. Phys. A*, **183**: 257 (1972)
- 62 M. Neffgen, G. Baldsieien, S. Frauendorf et al, *Nucl. Phys. A*, **595**: 499 (1995)
- 63 A. Afanasjev, D. Fossan, G. Lane, and I. Ragnarsson, *Phys. Rep.*, **322**: 1 (1999)
- 64 S. Frauendorf and J. Meng, *Nucl. Phys. A*, **617**: 131 (1997)
- 65 I. Kuti, Q. B. Chen, J. Timár et al, *Phys. Rev. Lett.*, **113**: 032501 (2014)
- 66 D. Jerrestam, B. Fogelberg, R. A. Bark et al, *Phys. Rev. C*, **52**: 2448 (1995)
- 67 J. Timár, T. Koike, N. Pietralla et al, *Phys. Rev. C*, **76**: 024307 (2007)
- 68 S. Sihotra, Z. Naik, R. Palit et al, *Eur. Phys. J. A*, **43**: 45 (2009)
- 69 P. Datta, S. Chattopadhyay, P. Banerjee et al, *Phys. Rev. C*, **67**: 014325 (2003)
- 70 C. Y. He, L. H. Zhu, X. G. Wu et al, *Phys. Rev. C*, **81**: 057301 (2010)
- 71 P. Ring, *Prog. Part. Nucl. Phys.*, **37**: 193 (1996)
- 72 D. Vretenar, A. V. Afanasjev, G. A. Lalazissis et al, *Phys. Rep.*, **409**: 101 (2005)
- 73 J. Meng, H. Toki, S. G. Zhou et al, *Prog. Part. Nucl. Phys.*, **57**: 470 (2006)
- 74 W. Koepf, and P. Ring, *Nucl. Phys. A*, **493**: 61 (1989)
- 75 H. Madokoro, J. Meng, M. Matsuzaki et al, *Phys. Rev. C*, **62**: 061301(R) (2000)
- 76 P. W. Zhao, *Phys. Lett. B*, **773**: 1 (2017)
- 77 P. W. Zhao, Y. K. Wang, and Q. B. Chen, *Phys. Rev. C*, **99**: 054319 (2019)
- 78 J. Peng, J. Meng, P. Ring, and S. Q. Zhang, *Phys. Rev. C*, **78**: 024313 (2008)
- 79 P. W. Zhao, S. Q. Zhang, J. Peng et al, *Phys. Lett. B*, **699**: 181 (2011)
- 80 P. W. Zhao, J. Peng, H. Z. Liang et al, *Phys. Rev. C*, **85**: 054310 (2012)
- 81 P. W. Zhao, J. Peng, H. Z. Liang et al, *Phys. Rev. Lett.*, **107**: 122501 (2011)
- 82 X. W. Li, J. Li, J. B. Lu et al, *Phys. Rev. C*, **86**: 057305 (2012)
- 83 K. Y. Ma, J. B. Lu, D. Yang et al, *Eur. Phys. J. A*, **48**: 82 (2012)
- 84 P. Zhang, B. Qi, and S. Y. Wang, *Phys. Rev. C*, **89**: 047302 (2014)
- 85 J. Peng and P. W. Zhao, *Phys. Rev. C*, **91**: 044329 (2015)
- 86 W. J. Sun, H. D. Xu, J. Li et al, *Chin. Phys. C*, **40**: 084101 (2016)
- 87 P. W. Zhao, Z. P. Li, J. M. Yao et al, *Phys. Rev. C*, **82**: 054319 (2010)
- 88 S. Frauendorf and J. Meng, *Z. Phys. A*, **356**: 263 (1996)