New Geiger-Nuttall law for cluster radioactivity half-lives^{*}

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Abstract: In this study, derived from Balasubramaniam's formula [Phys. Rev. C **70**, 017301 (2004)] and further considering the effect of the parent nucleus mass, blocking effect, and effect of reduced mass on cluster radioactivity half-lives, we propose a new Geiger-Nuttall law that is model-independent to systematically evaluate the half-lives of this process for 16 even-even nuclei and 10 odd-*A* nuclei. For comparison, a single universal curve for cluster radioactivity and α decay proposed by Poenaru [Phys. Rev. C **83**, 014601 (2011)], a scaling law proposed by Horoi [J. Phys. G: Nucl. Part. Phys. **30**, 945 (2004)], an extension of the Viola-Seaborg formula from α decay to cluster radioactivity proposed by Ren *et al.* [Phys. Rev. C **70**, 034304 (2004)], a new semi-empirical formula for exotic cluster decay proposed by Balasubramaniam *et al.* [Phys. Rev. C **70**, 017301 (2004)], and a unified formula for the half-lives of α decay and cluster radioactivity proposed by Ni *et al.* [Phys. Rev. C **78**, 044310 (2008)] are also used. The calculated results of our new Geiger-Nuttall law are in good agreement with the experimental half-lives, with the least rms being 0.606, and are better than the compared values. Moreover, we extend this formula to predict the cluster radioactivity half-lives of 51 nuclei whose decay energies are energetically allowed or observed but not yet quantified in NUBASE2020.

Keywords: cluster radioactivity, half-lives, Geiger and Nuttall law, decay energy

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I. INTRODUCTION

Spontaneous nuclear radioactivity embracing α decay, proton radioactivity, and cluster radioactivity has constantly been regarded as a vital channel to explore the nuclear structures of heavy and superheavy nuclei [1–8]. Cluster radioactivity, as one of the most important decay modes [9, 10], has been a prominent topic [11] since its first theoretical prediction was proposed by Săndulescu, Poenaru, and Greiner in 1980 [12] and experimental confirmation was realized from observing ¹⁴C emitted from ²²³Ra by Rose and Jones in 1984 [13]. This exotic decay mode, known as an intermediate physical process between α decay and spontaneous fission [14, 15], is gen-

erally dealt with via two main types of approaches: α -like and fission-like models [16–19]. For the former, the emitted cluster is considered pre-born in the parent nucleus before it penetrates the interacting potential barrier with different preformation probabilities [20–23]. For the latter, the emitted cluster is assumed to be formed along with the constant geometric shape deformations occurring in its nuclear barrier penetration process from the parent nucleus to the scission configuration within Gamow's penetration theory [24–26]. Essentially, the α like and fission-like models are equivalent because the cluster preformation probability in the α -like model is interpreted as the penetrability of the pre-scission part of the nuclear barrier in the fission-like model [27].

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Moreover, with the development of α decay and cluster radioactivity exploration, intensive phenomenological semi-empirical relationships based on the above theoretical approaches and/or models, as generalizations of the striking law of half-lives in logarithmic form and the decay energy in the case of α decay proposed by Geiger and Nuttall [28, 29], have been effectively applied to the investigation of cluster radioactivity [30-34] because they both share the quantum tunneling mechanism. For instance, in 2004. Ren et al. extended the famous Viola-Seaborg formula from α decay [35, 36] to complex cluster radioactivity (EVS) [37]. Later, derived from the WKB barrier penetration probability with several approximations, Ni et al. proposed a unified formula for the half-lives of both α decay and cluster radioactivity [38]. In 2009, using the microscopic mechanism of chargedparticle emission, Qi et al. proposed a universal decay law for charged-particle radioactivity including α decay and cluster radioactivity based on α -like *R*-matrix theory [39]. In 2011, Poenaru proposed a single universal curve based on fission-like theory for α decay and cluster radioactivity (UNIV) [40]. These formulas successfully contribute to describing cluster radioactivity half-lives.

Over the past years, systematic theoretical reinvestigation has been conducted on α decay owing to special interest in studies on superheavy nuclei and cluster radioactivity [41, 42]. The advancing progress of experimental techniques and infrastructures has enabled an increasing number of cluster members ranging from ¹⁴C to ³⁴Si, decaying from 221 Fr to 242 Cm with the daughter nuclei usually being the doubly magic nucleus ²⁰⁸Pb or its neighboring nucleus [43, 44], to be availably observed in experiments [45, 46]. Simultaneously, cluster radioactivity experimental data have been accumulated with adequately accurate precision so that their systematics can be explored almost independently of a model. Nevertheless, all analyses or predictions of cluster radioactivity half-lives undertaken to date were mostly dependent on theoretical models. The prescriptions of these models and/or the physical meaning of the parameters used by them have been subjected to criticism to various degrees [47].

Consequently, a number of model-independent semiempirical formulas were established to evaluate the halflives of cluster radioactivity, which were directly derived from experimental facts dependent on cluster radioactivity half-lives. In 2004, Horoi *et al.* proposed a scaling law underlining the most important variables that scale the cluster radioactivity probabilities [47]. Before long, Balasubramaniam and his collaborators put forward a semi-empirical formula based on experimental facts, which had only three parameters with a root-mean-square (rms) deviation $\sigma = 0.89$ [48]. However, there are still large deviations between experimental cluster radioactivity half-lives and the theoretically calculated values obtained using these formulas. It is necessary to probe more new physical factors critical to cluster radioactivity halflives, such as the parent nucleus mass, to improve the accuracy. To this end, in the present study, we systematically analyze the effect of the parent mass on cluster radioactivity half-lives and find that there is a linear relationship between cluster radioactivity half-lives and the parent nucleus mass. Furthermore, based on Balasubramaniam's work, considering the effect of the parent mass nucleus, blocking effect, and effect of reduced mass on cluster radioactivity half-lives, we propose a new Geiger-Nuttall (G-N) law that is model-independent to calculate cluster radioactivity half-lives.

This article is organized as follows. The detailed theoretical framework of the phenomenological semi-empirical formula of cluster radioactivity half-lives is presented in Sec. II. Numerical results and detailed discussion are given in Sec. III. Finally, Sec. IV contains a brief summary.

II. THEORETICAL FRAMEWORK

In 1911, Geiger and Nuttall first observed the linear relationship between the α decay half-life and *Q*-value and proposed the G-N law, which can be expressed as [49]

$$\log_{10}(T_{1/2}) = aQ^{-1/2} + b, \tag{1}$$

where *a* and *b* are two adjustable parameters dependent on different α decay radioactive isotopes, which makes it difficult to universally describe all α decay cases using this formula and results in a limited predictive power. In 2000, considering the effect of the parent nucleus mass and charge on α decay half-lives, Royer generalized the G-N law to estimate the half-lives of all α decay cases in a uniform form. This can be written as [50]

$$\log_{10}(T_{1/2}) = a \frac{Z}{\sqrt{Q_{\alpha}}} + bA^{1/6}Z^{1/2} + c,$$
 (2)

where Z and A are the charge and mass number of the parent nucleus, Q_{α} is the α decay energy, and a, b, and c are the adjustable parameters for different types of parent nuclei. The corresponding values of these parameters are given as follows:

Case I for even-even nuclei: a=-25.31, b=-1.1629, c=1.5864;

Case II for even-odd nuclei: *a*=-26.65, *b*=-1.0859, *c*=1.5848;

Case III for odd-even nuclei: a=-25.68, b=-1.1423,

c=1.592;

Case IV for odd-odd nuclei: *a*=-29.48, *b*=-1.113, *c*=1.6971.

However, several physical factors in contrast with α decay should be considered differently when generalizing the G-N law to describe the half-lives of cluster radioactivity in spite of the two similar physical processes. For instance, in 2004, Horoi proposed the first model-independent scaling law for the half-lives of α decay and cluster radioactivity [47], which can be expressed as

$$\log_{10}(T_{1/2}) = (a_1\mu^x + b_1) \left[\frac{(Z_e Z_d)^y}{\sqrt{Q}} - 7 \right] + (a_2\mu^x + b_2), \quad (3)$$

where Z_e and Z_d are the charge number of the emitted α particle and/or cluster and daughter nucleus, respectively, $a_1=9.1$, $b_1=-10.2$, $a_2=7.39$, $b_2=-23.2$, x=0.416, and y=0.613 are the parameters adjusted by fitting the experimental data of 119 α decays and 11 cluster radioactivities emitted from even-even parent nuclei, and Q is the decay energy of α decay and/or cluster radioactivity. In the study, the effect of the reduced mass μ on both α decay and cluster radioactivity half-lives was systematically investigated. The results indicated that μ is sensitive to the half-lives of cluster radioactivity, whereas it has less effect on α decay half-lives.

In addition, a recent study reported that cluster radioactivity half-lives increase as the size of the emitted cluster increases [32, 38]. Therefore, the mass number (A_c) and charge number (Z_c) of the emitted cluster are crucial factors in the investigation of cluster radioactivity half-lives. Meanwhile, because the same cluster is emitted from different parent nuclei, the mass asymmetry $(\eta = \frac{A_d - A_c}{A})$, where A_d is the mass number of the daughter nucleus) and charge asymmetry $(\eta_z = \frac{Z_d - Z_c}{Z})$, where Z_c is the charge number of the emitted cluster) should be further explored. In 2004, considering the above two experimental factors, Balasubramaniam *et al.* put forward a new model-independent formula, which is expressed as [48]

$$\log_{10}(T_{1/2}) = \frac{aA_c\eta + bZ_c\eta_z}{\sqrt{Q_c}} + c,$$
 (4)

where Q_c is the cluster radioactivity decay energy, and a=10.603, b=78.027, and c=-80.669 are the adjustable parameters.

Generally, the structures of the ground states of odd-*A* and odd-odd nuclei are naturally different from those of even-even nuclei because of the unpaired nucleon which hinders the transition of these states, a process known as

the blocking effect associated with the preformation probability [37]. Moreover, the larger this hindrance, the smaller the preformation probability, and a smaller preformation probability results in a prolonged cluster radioactivity half-life [28]. Therefore, the blocking effect is another critical factor in the calculation of cluster radioactivity half-lives. Based on the above discussion, considering the effect of the parent nucleus mass, effect of mass and charge asymmetries, blocking effect, and effect of reduced mass on cluster radioactivity half-lives, we propose a new G-N law for cluster radioactivity half-lives.

III. RESULTS AND DISCUSSION

In this study, we propose a new G-N law to systematically study the cluster radioactivity half-lives of 26 nuclei. To intuitively validate the applicability of the G-N law to cluster radioactivity, the dependence between the experimental cluster radioactivity half-lives in logarithmic form and $Q_c^{-1/2}$ for five cluster radioactive isotope cases is plotted in Fig. 1. The cluster radioactivity decay energy Q_c is calculated as below.

$$Q_{c} = B(A_{c}, Z_{c}) + B(A_{d}, Z_{d}) - B(A, Z).$$
 (5)

Here, $B(A_c, Z_c)$, $B(A_d, Z_d)$, and B(A, Z) are the binding energy of the emitted cluster, daughter, and parent nucleus, respectively. They are all taken from the latest 2020 atomic mass table AME2020 [51] and NUBASE2020 [52]. From Fig. 1, we can see that the experimental cluster radioactivity half-life in logarithmic form exhibits a linear relationship with $Q_c^{-1/2}$ for each case. Therefore, the construction of a semi-empirical formula for cluster radioactivity half-lives based on the G-N law is both physically reasonable and analytically numerical. Other than the cluster radioactive isotopes plotted above, there



Fig. 1. (color online) G-N law plots of cluster radioactivity for different isotopes.

are still several individual cluster radioactivity cases with available experimental data that contribute to the study of cluster radioactivity, which should also be investigated. To break the isotopic dependent limit of the G-N law and further generalize it to all cluster radioactivity cases, it is necessary to probe more physical factors behind this phenomenon and make several advancements on the original G-N law.

Experimentally, in the case of α decay, the reduced mass changes by 0.3% as the mass number of the daughter nucleus (A_d) changes from 200 to 300 [38]. It is rational to consider the reduced mass as a constant in the calculations of the half-lives of α decay. Nevertheless, the cluster radioactivity half-lives are extremely dependent on the reduced mass because its value promptly varies to both fragments of the emitted cluster and daughter nucleus [47]. In reality, for 14 C radioactivity, the reduced mass changes by 0.15% as the mass number of the daughter nucleus (A_d) changes from 207 to 212. As for ²⁴Ne radioactivity, the reduced mass changes by 3.8% as the mass number of the daughter nucleus (A_d) changes from 206 to 210. Moreover, Horoi also reported a proportional relationship between cluster radioactivity half-lives and the square root of the reduced mass $\sqrt{\mu}$ [47]. Balasubramaniam's work has shown that cluster radioactivity half-lives in logarithmic form are dependent on both mass asymmetry η and charge asymmetry η_z [48]. Furthermore, the correlation between η and η_z is known to be weak, and they should be quantified separately [53, 54]. Therefore, it is interesting to explore whether the product of the square root of the reduced mass $\sqrt{\mu}$ and mass asymmetry η as well as the product of the square root of the reduced mass $\sqrt{\mu}$ and charge asymmetry η_z can be effectively supplemented into Balasubramaniam's original formula to improve the accuracy. For a distinct insight into the correlation between cluster radioactivity half-lives in logarithmic form and $\sqrt{\mu}\eta$ and $\sqrt{\mu}\eta_z$, the correlations in the cases of ¹⁴C emitted from Ra isotopes and ²⁴Ne emitted from U isotopes are plotted in Fig. 2 and Fig. 3, where the x axes represent $\sqrt{\mu}\eta$ and/or $\sqrt{\mu}\eta_z$, respectively, and the y axes denote the cluster radioactivity halflives in logarithmic form. From these two figures, it is clear that the cluster radioactivity half-lives in logarithmic form are linear to both $\sqrt{\mu\eta}$ and $\sqrt{\mu\eta_z}$. Then, Balasubramaniam's formula can be revised as

$$\log_{10}(T_{1/2}) = \frac{a_1 A_c \eta + a_2 Z_c \eta_z}{\sqrt{Q_c}} + \sqrt{\mu} (a_3 \eta + a_4 \eta_z) + a_5.$$
(6)

Moreover, because the $a_3 \sqrt{\mu}\eta$ and $a_4 \sqrt{\mu}\eta_z$ terms separately reflect the information of both η and η_z , to minimize the number of adjustable parameters, we make attempts to combine the first term $a_1 \frac{A_c \eta}{\sqrt{Q_c}}$ and second term



Fig. 2. (color online) Correlation between the experimental cluster radioactivity half-lives in logarithmic form and $\sqrt{\mu}\eta$ in the cases of ¹⁴C emitted from Ra isotopes and ²⁴Ne emitted from U isotopes.



Fig. 3. (color online) Correlation between the experimental cluster radioactivity half-lives in logarithmic form and $\sqrt{\mu}\eta_z$ in the cases of ¹⁴C emitted from Ra isotopes and ²⁴Ne emitted from U isotopes.

 $a_2 \frac{Z_c \eta_z}{\sqrt{Q_c}}$ as one term $a_1 \frac{A_c \eta + Z_c \eta_z}{\sqrt{Q_c}}$. To justify the feasibility of this alteration, we plot the relationship between this term and the experimental half-lives of cluster radioactivity in logarithmic form in Fig. 4. This figure distinctly demonstrates that the experimental half-lives of cluster radioactivity in logarithmic form lie almost in a straight line against this combined term, as expected. Therefore, by further considering the blocking effect of odd-*A* nuclei characterized by the blocking factor *h*, Eq. 6 can be expressed as

$$\log_{10}(T_{1/2}) = a_1 \frac{A_c \eta + Z_c \eta_z}{\sqrt{Q_c}} + \sqrt{\mu}(a_2 \eta + a_3 \eta_z) + a_4 + h, \quad (7)$$



Fig. 4. (color online) Correlation between the experimental cluster radioactivity half-lives in logarithmic form and the quantity $\frac{A_c \eta + Z_c \eta_z}{\sqrt{Q_c}}$.

where a_1 , a_2 , a_3 , and a_4 are the adjustable parameters, and σ represents the standard deviation between the experimental cluster radioactivity half-lives and the calculated values. In this study, it is defined as

$$\sigma = \sqrt{\sum (\log_{10} T_{1/2}^{\exp} - \log_{10} T_{1/2}^{cal})^2 / n},$$
(8)

where $\log_{10}T_{1/2}^{\exp}$ and $\log_{10}T_{1/2}^{cal}$ are the experimental cluster radioactivity half-lives in logarithmic form and calculated values, respectively, and *n* is the number of nuclei involved for each case. Based on Eq. (7), the values of the adjustable parameters obtained using the least squares approach to fit the available experimental cluster radioactivity half-lives of 16 even-even emitters are a_1 = 31.081881, a_2 =24.938505, a_3 =-30.311899, and a_4 = -66.443373.

The blocking factor h=1.034577 is determined by fitting the experimental cluster radioactivity half-lives of 10 odd-*A* nuclei while a_1 , a_2 , a_3 , and a_4 are fixed. To display the correlation between the quantity $(\log_{10}T_{1/2}^{\exp} - \sqrt{\mu}(a_2\eta + a_3\eta_z) - a_4 - h_{\log})$ and the term $a_1 \frac{A_c\eta + Z_c\eta_z}{\sqrt{Q_c}}$, we plot this correlation using the set of adjustable parameters between the quantity $(\log_{10}T_{1/2}^{\exp} - \sqrt{\mu}(a_2\eta + a_3\eta_z) - a_4 - h_{\log})$ and the term $a_1 \frac{A_c\eta + Z_c\eta_z}{\sqrt{Q_c}}$ shown in Fig. 5. As clearly shown in this figure, there is a linear relationship between the experimental cluster radioactivity half-lives in logarithmic form and the results calculated using Eq. (7). The standard deviation between the experimental half-lives and calculated results for the 26 emitters is $\sigma=0.739$, which still implies large deviations between the experimental cluster radioactivity half-lives and calcu-



Fig. 5. (color online) Linear correlation between the quantity $[\log_{10}T_{1/2}^{exp} - \sqrt{\mu}(a_2\eta + a_3\eta_z) - a_4 - h]$ and the term $a_1 \frac{A_c\eta + Z_c\eta_z}{\sqrt{Q_c}}$ of Eq. (7).

lated values.

Previous phenomenological semi-empirical relationships concerning cluster radioactivity including Eq. (7) have placed considerable emphasis on the significance of the emitted cluster, which may impact the calculation of cluster radioactivity half-lives. In 2000, Royer indicated that the mass of the parent nucleus A is a key component in the calculation of α decay half-lives. With the similarity in the physical processes of α decay and cluster radioactivity, it is interesting to explore whether the mass of the parent nucleus A may also have an impact on cluster radioactivity half-lives. To this end, we plot the behavior of the experimental cluster radioactivity half-lives in logarithmic form against the mass of the parent nucleus A in-Fig. 6. Surprisingly, there is a visibly linear dependence between the experimental cluster radioactivity half-lives in logarithmic form and the mass of the parent nucleus A. The results illustrated in this figure verify the significance of probing the mass of the parent nucleus when estimating cluster radioactivity half-lives. Finally, based on Eq. (7), by further considering the effect of the parent nucleus mass A on cluster radioactivity half-life, we propose a new G-N law to evaluate cluster radioactivity halflives, which can be expressed as

$$\log_{10}(T_{1/2}) = a \frac{A_c \eta + Z_c \eta_z}{\sqrt{Q_c}} + \sqrt{\mu} (b\eta + c\eta_z) + dA + e + h_{\log}.$$
 (9)

Here *a*, *b*, *c*, *d*, and *e* are the adjustable parameters, and h_{\log} is the blocking factor. By fitting the experimental half-lives of cluster radioactivity for the 16 even-even emitters while $h_{\log}=0$, we obtain the values of these adjustable parameters as a=32.500137, b=18.117466, c=-29.999023, d=0.272744, and e=-111.134251.



Fig. 6. (color online) Correlation between the experimental cluster radioactivity half-lives in logarithmic form and the parent nucleus mass *A*.

Moreover, $h_{log}=1.290933$ is obtained by fitting the experimental half-lives of cluster radioactivity for 10 odd-A nuclei while a, b, c, d, and e are fixed. Using this set of parameters, the correlation between the quantity $\left[\log_{10}T_{1/2}^{\exp} - \sqrt{\mu}(b\eta + c\eta_z) - dA - e - h_{\log}\right]$ and the term $a\frac{A_c\eta + Z_c\eta_z}{\sqrt{2}}$ is plotted in Fig. 7. From this figure, it is clear that the experimental cluster radioactivity half-lives in logarithmic form are proportional to the calculated results using our new G-N law. The standard deviation between the experimental half-lives and the calculated values obtained using Eq. (9) for the 26 emitters is 0.606. Compared to the calculated results obtained using Eq. (7), the standard deviation σ dramatically decreases by 0.739 - 0.606=18.00%. Furthermore, the difference 0.739 between the experimental half-lives and the calculated values using Eq. (7) and our new G-N law is plotted in Fig. 8. From this figure, we find that the distribution of the discrepancy of Eq. (7) is slightly scattered compared to that of the new G-N law, suggesting that the mass of the parent nucleus A is another important physical factor to be considered for cluster radioactivity half-lives, which may have been previously ignored.

For comparison, the cluster radioactivity half-lives are also evaluated using UNIV [40], Ni's formula [38], EVS [37], the scaling law [47], and Balasubramaniam's formula [48]. The detailed calculated results are given in Table 1. In this table, the first two columns denote the decay process and cluster radioactivity decay energy, which are denoted as Decay and Q_c , respectively. The experimental cluster radioactivity half-lives in logarithmic form extracted from Ref. [31], Ref. [44], and Ref. [55] are listed in the third column, denoted as $\lg T_{1/2}^{\exp}$. The fourth to tenth columns contain the calculated results obtained using UNIV, Ni's formula, the scaling law, Balasub-



Fig. 7. (color online) Linear correlation between the quantity $[\log_{10}T_{1/2}^{\exp} - \sqrt{\mu}(b\eta + c\eta_z) - dA - e - h_{\log}]$ and the term $a\frac{A_c\eta + Z_c\eta_z}{\sqrt{Q_c}}$ of Eq. (9).



Fig. 8. (color online) Differences between the experimental cluster radioactivity half-lives in logarithmic form and the calculated values obtained using Eqs. (7) and (9).

ramaniam's formula, EVS, Eq. (7), and our new G-N law, which are labeled as $\lg T_{1/2}^{\text{UNIV}}$, $\lg T_{1/2}^{\text{Ni}}$, $\lg T_{1/2}^{\text{SL}}$, $\lg T_{1/2}^{\text{Bala}}$, $\lg T_{1/2}^{\text{EVS}}$, $\lg T_{1/2}^{\text{Cal}^1}$, and $\lg T_{1/2}^{\text{Cal}^2}$, respectively. In this table, it is clear that the results calculated using EVS, Eq. (7), and our new G-N law are better than those calculated using UNIV, the scaling law, Balasubramaniam's formula, and Ni's formula.

Furthermore, the calculated results of σ using UNIV, Ni's formula, the scaling law, EVS, Balasubramaniam's formula, Eq. (7), and our new G-N law are also provided in Table 2. From this table, it is clear that for even-even nuclei, the values of σ obtained using EVS and our new G-N law are essentially consistent and smaller than those

 Table 1.
 Comparisons between the experimental cluster radioactivity half-lives (in seconds) in logarithmic form and the calculated values using different formulas.

Decay	$Q_{\rm c}/{ m MeV}$	$\lg T_{1/2}^{\exp}$	$\lg T_{1/2}^{\mathrm{UNIV}}$	$\lg T_{1/2}^{\rm Ni}$	$\lg T_{1/2}^{\mathrm{SL}}$	$\log T_{1/2}^{\text{Bala}}$	$\log T_{1/2}^{\rm EVS}$	$\lg T_{1/2}^{\operatorname{Cal}^1}$	$\lg T_{1/2}^{\operatorname{Cal}^2}$
221 Fr \rightarrow 207 Tl $+$ 14 C	$31.29^{+0.00707}_{-0.00707}$	14.56	14.22	14.63	13.54	14.65	14.43	15.51	$15.17^{+0.01142}_{-0.01142}$
221 Ra \rightarrow ²⁰⁷ Pb+ 14 C	$32.40^{+0.00512}_{-0.00512}$	13.39	13.12	13.48	12.27	13.14	13.44	13.72	$13.31^{+0.00786}_{-0.00785}$
222 Ra \rightarrow 208 Pb+ 14 C	$33.05^{+0.00415}_{-0.00415}$	11.22	11.95	11.02	11.00	12.23	10.74	11.84	$11.38^{+0.00618}_{-0.00618}$
223 Ra \rightarrow 209 Pb+ 14 C	$31.83^{+0.00270}_{-0.00270}$	15.05	14.04	14.56	13.42	14.01	14.61	14.75	$14.89^{+0.00426}_{-0.00426}$
224 Ra \rightarrow 210 Pb+ 14 C	$30.53^{+0.00228}_{-0.00228}$	15.87	16.42	15.86	16.14	16.01	15.98	15.82	$16.06^{+0.00383}_{-0.00383}$
226 Ra \rightarrow 212 Pb $+$ 14 C	$28.20\substack{+0.00262\\-0.00262}$	21.20	21.20	20.94	21.53	19.97	21.47	19.99	$20.92\substack{+0.00496\\-0.00496}$
$^{223}\text{Ac} \rightarrow ^{209}\text{Bi} + {}^{14}\text{C}$	$33.06^{+0.00714}_{-0.00714}$	12.60	12.73	13.19	11.93	12.35	13.37	12.82	$12.89^{+0.01064}_{-0.01063}$
$^{225}\text{Ac} \rightarrow ^{211}\text{Bi} + {}^{14}\text{C}$	$30.48^{+0.00707}_{-0.00707}$	17.16	17.40	18.24	17.26	16.24	18.84	16.93	$17.68\substack{+0.01192\\-0.01191}$
228 Th \rightarrow 208 Pb $+$ 20 O	$44.72^{+0.00229}_{-0.00229}$	20.73	21.90	21.54	21.20	22.22	20.97	22.15	$21.61\substack{+0.00287\\-0.00287}$
231 Pa \rightarrow 208 Pb $+$ 23 F	$51.88^{+0.03007}_{-0.03007}$	26.02	24.47	25.59	23.78	24.65	24.62	25.46	$25.36\substack{+0.03354\\-0.03351}$
230 Th \rightarrow ²⁰⁶ Hg+ ²⁴ Ne	$57.76^{+0.02004}_{-0.02004}$	24.63	24.68	24.58	23.92	25.68	24.17	25.22	$24.36\substack{+0.01987\\-0.01986}$
231 Pa \rightarrow 207 Tl+ 24 Ne	$60.41\substack{+0.00534\\-0.00534}$	22.89	22.15	23.09	21.32	23.59	23.44	23.77	$23.30\substack{+0.00495\\-0.00495}$
$^{232}U \rightarrow ^{208}Pb + ^{24}Ne$	$62.31^{+0.00217}_{-0.00217}$	20.39	20.76	20.36	19.94	22.26	21.00	21.03	$20.48\substack{+0.00192\\-0.00192}$
$^{233}U \rightarrow ^{209}Pb + ^{24}Ne$	$60.49^{+0.00290}_{-0.00290}$	24.84	23.11	24.41	22.55	23.83	24.76	23.83	$23.86\substack{+0.00270\\-0.00270}$
$^{234}U \rightarrow ^{210}Pb + ^{24}Ne$	$58.82^{+0.00185}_{-0.00185}$	25.93	25.38	25.81	25.03	25.32	25.61	24.48	$24.56\substack{+0.00179\\-0.00179}$
$^{235}\text{U} \rightarrow ^{211}\text{Pb} + ^{24}\text{Ne}$	$57.36^{+0.00260}_{-0.00260}$	27.42	27.47	29.51	27.31	26.69	29.06	27.08	$27.72^{+0.00262}_{-0.00262}$
$^{233}U \rightarrow ^{208}Pb + ^{25}Ne$	$60.70\substack{+0.02911\\-0.02911}$	24.84	23.75	24.88	23.05	24.43	24.48	24.61	$24.57\substack{+0.02748\\-0.02746}$
$^{234}U \rightarrow ^{208}Pb + ^{26}Ne$	$59.41^{+0.01807}_{-0.01807}$	25.93	26.45	26.52	25.84	26.37	24.80	25.86	$25.83^{+0.01799}_{-0.01798}$
$^{234}U \rightarrow ^{206}Hg + ^{28}Mg$	$74.1\substack{+0.02003\\-0.02003}$	25.53	25.19	25.25	24.76	25.95	25.11	25.43	$24.92^{+0.01540}_{-0.01539}$
$^{236}U \rightarrow ^{208}Hg + ^{28}Mg$	$70.73^{+0.03002}_{-0.03002}$	27.58	29.37	30.33	29.28	28.54	29.11	28.45	$28.53\substack{+0.02479\\-0.02478}$
236 Pu \rightarrow 208 Pb $+$ 28 Mg	$79.67\substack{+0.00213\\-0.00213}$	21.52	21.19	20.76	20.83	22.82	21.90	21.40	$21.20\substack{+0.00147\\-0.00147}$
238 Pu \rightarrow 210 Pb $+$ 28 Mg	$75.91^{+0.00180}_{-0.00180}$	25.70	25.35	25.96	25.42	25.42	26.00	24.41	$24.81\substack{+0.00134\\-0.00134}$
$^{236}U \rightarrow ^{206}Hg + ^{30}Mg$	$72.27\substack{+0.02007\\-0.02007}$	27.58	29.10	29.47	28.69	28.64	27.25	28.31	$28.35\substack{+0.01659\\-0.01658}$
238 Pu \rightarrow 208 Pb $+$ 30 Mg	$76.80\substack{+0.00203\\-0.00203}$	25.70	25.97	26.10	25.71	26.05	25.01	24.76	$25.13\substack{+0.00154\\-0.00154}$
238 Pu \rightarrow 206 Hg+ 32 Si	$91.19\substack{+0.02003 \\ -0.02003}$	25.28	25.50	25.59	25.70	25.63	25.34	25.65	$25.71^{+0.01242}_{-0.01242}$
$^{242}Cm \rightarrow ^{208}Pb + ^{34}Si$	$96.51^{+0.00175}_{-0.00175}$	23.15	23.91	23.48	24.21	24.46	23.16	23.12	$24.08^{+0.00103}_{-0.00103}$

Table 2. Standard deviation σ between the experimental data and the calculated values using different formulas for cluster radioactivity.

σ	UNIV	Ni	SL	Bala	EVS	Cal ¹	Cal ²
even-even (n=16)	0.769	0.913	0.731	0.997	0.595	0.819	0.662
odd-A (n=10)	0.918	0.808	1.462	0.784	0.935	0.590	0.499
total (<i>n</i> =26)	0.829	0.874	1.073	0.921	0.744	0.738	0.606

calculated using UNIV, Ni's formula, the scaling law, Balasubramaniam's formula, and Eq. (7). For odd-*A* nuclei, the value of σ calculated using our new G-N law is $\sigma = 0.499$, which is the smallest out of all the results. Moreover, for all 26 nuclei, the value of σ calculated using our new G-N law is $\sigma = 0.606$, corresponding to a factor of 4 on average, which is comparably less than the results of all compared formulas. By comparing the results of EVS, which are the closest to our results from the new G-N law, the standard deviation of the half-lives for 0.744 - 0.606 = 18.55%. Com-26 nuclei decreases by 0.744 pared to the results from Balasubramaniam's original ver-G-N improves sion, our new law by $\frac{0.921 - 0.606}{2} = 34.20\%$. This means that the results calcu-0.921 lated using our new G-N law are in good agreement with the experimental cluster radioactivity half-lives. It should be noted that the differences in physical object selection and the corresponding construction conception using different numbers of parameters for different analytical for-Balasubr

mulas cause the differences in rms deviation. To make further intuitive comparisons, we plot the discrepancies between the experimental cluster radioactivity half-lives in logarithmic form and the calculated values obtained using UNIV, Ni's formula, the scaling law, EVS, Balasubramaniam's formula, Eq. (7), and our new G-N law in Fig. 9. In this figure, the discrepancies are depicted in two parts: comparisons to model-dependent formulas (UNIV, Ni's formula, and EVS) and modelindependent formulas (scaling law and Balasubramaniam's formula). We can distinctly see that compared with the model-dependent formulas, the distributions of the discrepancies for EVS, Eq. (7), and our new G-N law are more converging on the zero area based on the overall trend. The calculated results of our new G-N law are within ± 1 overall, except for the case of $^{234}\text{U} \rightarrow ^{210}\text{Pb} + ^{24}\text{Ne}$, which has a discrepancy of 1.37, perhaps owing to the imprecision of measurement. For the calculated results of EVS and Eq. (7), both have five cases whose discrepancies are outside the scale of ± 1 for all 26 nuclei. For the discrepancies of UNIV, there are nine cases beyond of the scale of ± 1 . For the calculated results of Ni's formula, the case of $^{235}U \rightarrow ^{211}Pb + ^{24}Ne$, which has a discrepancy of 2.089, and the case of 208×10^{-11} km s²³⁶ m s²³⁶ m s²³⁶ m s²⁸ ${}^{6}Pu \rightarrow {}^{208}Pb + {}^{28}Mg$, which exhibits a discrepancy of



Fig. 9. (color online) Comparison of the discrepancies between the experimental cluster radioactivity half-lives in logarithmic form and the calculated values obtained using different semi-empirical formulas and our new G-N law. The top diagram presents the comparison with the calculated results obtained using the model-independent semi-empirical formulas of the scaling law and Balasubramaniam's formula. The bottom diagram presents the comparison with the calculated results obtained using the model-dependent semi-empirical formulas of UNIV, Ni's formula, and EVS.

2.748, are outside the scale of ± 2 . For the comparisons with model-independent formulas, the discrepancies of Balasubramaniam's formula are within ± 2 overall, with 11 nuclei whose discrepancies are beyond of the scale of ± 1 . For the scaling law, there are 10 nuclei whose discrepancy of 2.242 in the case of $^{231}Pa \rightarrow ^{208}Pb + ^{23}F$ and a discrepancy of 2.290 in the case of $^{233}U \rightarrow ^{209}Pb + ^{24}Ne$. The results demonstrated in Fig. 9 further show that our new G-N law can reproduce the experimental cluster radio-activity half-lives well.

Furthermore, the cluster radioactivity half-lives are extremely sensitive to Q_c values. A change of 1 MeV in Q_c can lead to an uncertainty ranging from 10^2 to 10^3 times on the cluster radioactivity half-life [56]. Because the experimental values of binding energies have uncertainties, the theoretical values of Q_c and cluster radioactivity half-lives also have uncertainties. However, the uncertainties of Q_c cannot be directly experimentally measured. The uncertainties of Q_c can possibly be evaluated through the error propagation formula. If there is a linear correlation between the physical quantity whose uncertainty needs to be estimated and other quantities whose uncertainties have been experimentally measured,

$$y = a_0 + a_{x_1}x_1 + a_{x_2}x_2 + a_{x_3}x_3, \tag{10}$$

where a_{x_1} , a_{x_2} , and a_{x_3} are the coefficients of x_1 , x_2 , and x_3 , respectively, and a_0 is simply a constant. The uncertainties of the physical quantity can be indirectly calculated using the following formula:

$$\sigma^{2} = a_{x_{1}}^{2} \sigma_{x_{1}}^{2} + a_{x_{2}}^{2} \sigma_{x_{2}}^{2} + a_{x_{3}}^{2} \sigma_{x_{3}}^{2} + 2a_{x_{1}}a_{x_{2}} \operatorname{Cov}(x_{1}, x_{2}) + 2a_{x_{1}}a_{x_{3}} \operatorname{Cov}(x_{1}, x_{3}) + 2a_{x_{2}}a_{x_{3}} \operatorname{Cov}(x_{2}, x_{3}),$$
(11)

where σ_{x_1} , σ_{x_2} , and σ_{x_3} are the uncertainties of x_1 , x_2 , and x_3 , respectively, and $Cov(x_1, x_2)$, $Cov(x_1, x_3)$, and $Cov(x_2, x_3)$ denote the covariance between x_1 and x_2 , the covariance between x_1 and x_3 , and the covariance between x_2 and x_3 , respectively. If the variables x_1 , x_2 , and x_3 are independent, Eq. (11) can be simplified as

$$\sigma^2 = a_{x_1}^2 \sigma_{x_1}^2 + a_{x_2}^2 \sigma_{x_2}^2 + a_{x_3}^2 \sigma_{x_3}^2.$$
(12)

In the case of cluster radioactivity, the uncertainties of Q_c can be obtained using

$$\Delta Q_c = \sqrt{\Delta^2 B(A_c, Z_c) + \Delta^2 B(A_d, Z_d) + \Delta^2 B(A, Z)}, \quad (13)$$

where $\Delta B(A_c, Z_c)$, $\Delta B(A_d, Z_d)$, and $\Delta B(A, Z)$ represent the experimental uncertainties of $B(A_c, Z_c)$, $B(A_d, Z_d)$, and

	1 401						-EVS	
Decay	$Q_c/{ m MeV}$	$\lg T_{1/2}^{enp}$	$\lg T_{1/2}^{ONV}$	$\lg T_{1/2}^{\mathrm{INI}}$	$\lg T_{1/2}^{\operatorname{SL}}$	$\lg T_{1/2}^{\text{Bala}}$	$\lg T_{1/2}^{EV3}$	$\lg T_{1/2}^{\operatorname{Cal}}$
$^{219}\text{Rn} \rightarrow ^{205}\text{Hg} + {}^{14}\text{C}$	28.10	-	19.77	20.29	19.65	19.75	20.33	20.16
$Rn \rightarrow Hg^{+1}C$	28.54	_	18.78	18.05	18.61	18.99	17.86	18.40
$Fr \rightarrow {}^{200}Hg + {}^{10}N$	34.12	-	21.23	22.16	20.81	21.32	24.12	19.67
223 Ra $\rightarrow ^{203}$ Hg $+^{16}$ O	40.30	_	25.27	26.41	24.74	24.99	27.35	23.72
$^{223}\text{Ra} \rightarrow ^{211}\text{Pb} + ^{14}\text{C}$	29.47	-	18.51	19.37	18.53	17.76	19.81	19.54
$^{223}\text{Ra} \rightarrow ^{203}\text{Hg} + ^{20}\text{O}$	40.48	-	27.31	28.34	26.86	27.00	27.01	27.77
220 Ra $\rightarrow ^{200}$ Hg+ 20 O	40.82	_	26.62	26.41	26.18	26.59	24.98	26.40
$Ac \rightarrow {}^{208}Pb + {}^{13}N$	39.47	>14.76	13.84	14.47	12.90	14.50	16.53	12.74
$Ac \rightarrow {}^{20}Tl + {}^{20}O$	43.09	-	23.61	24.55	22.99	23.94	23.75	24.79
$Ac \rightarrow Hg^{200}Hg^{+25}F$	48.35	-	27.91	29.14	27.30	27.93	27.24	29.08
220 Th \rightarrow 208 Pb $+$ 18 O	45.73	>16.76	18.20	17.81	17.38	18.96	19.32	16.45
220 Th \rightarrow 212 Po+ 14 C	30.55	>15.36	18.13	17.83	18.10	16.27	18.61	16.51
227 Th \rightarrow 209 Pb $+$ 18 O	44.20	_	20.54	21.69	19.96	20.68	23.25	19.91
$^{228}\text{Th} \rightarrow ^{206}\text{Hg} + ^{22}\text{Ne}$	55.74	-	25.73	26.07	25.14	25.83	27.03	23.36
229 Th $\rightarrow 209$ Pb $+20$ O	43.40	_	24.08	25.25	23.64	23.80	24.62	24.99
$^{229}\text{Th} \rightarrow ^{203}\text{Hg} + ^{24}\text{Ne}$	57.83	_	24.65	25.72	23.80	25.59	25.48	25.16
231 Th $\rightarrow ^{207}$ Hg $+^{24}$ Ne	56.25	-	26.85	28.36	26.29	27.12	27.69	27.61
231 Th $\rightarrow ^{206}$ Hg $^{+23}$ Ne	56.80	-	27.02	28.30	26.31	27.42	26.92	28.04
232 Th $\rightarrow ^{208}$ Hg $+^{24}$ Ne	54.67	>29.2	29.26	29.86	28.89	28.71	28.62	28.42
232 Th $\rightarrow ^{200}$ Hg $^{+20}$ Ne	55.91	>29.2	29.31	29.45	28.66	29.10	26.79	29.06
227 Pa \rightarrow 209 Bi $+$ 18 O	45.87	-	18.96	20.01	18.27	19.00	21.86	17.79
229 Pa $\rightarrow ^{207}$ Tl 22 Ne	58.96	_	22.31	23.63	21.67	23.16	25.38	21.76
$^{230}U \rightarrow ^{208}Pb + ^{22}Ne$	61.39	>18.2	20.20	20.09	19.55	21.35	22.18	18.49
$U \rightarrow 200 Pb + 24 Ne$	61.35	>18.2	22.14	21.78	21.27	23.00	22.23	20.49
$U \rightarrow Hg^{205}Hg^{20}Mg$	74.32	>22.26	25.08	24.93	24.47	25.73	24.88	23.84
$U \rightarrow U^{203}Hg + U^{20}Mg$	74.23	>27.59	25.12	26.33	24.60	25.84	26.38	25.66
$U \rightarrow 210^{211} Pb + 24^{24} Ne$	57.36	>27.65	27.47	29.51	27.31	26.69	29.06	27.72
$U \rightarrow {}^{210}\text{Pb} + {}^{25}\text{Ne}$	57.68	>27.65	27.99	29.86	27.72	27.21	28.61	28.40
$U \rightarrow U^{207}Hg + {}^{20}Mg$	72.43	>28.45	27.22	29.00	26.98	27.22	28.47	27.99
$U \rightarrow Hg^{236}Hg^{-2}Mg$	72.48	>28.45	28.03	29.67	27.66	27.83	28.41	28.50
$U \rightarrow Pb+Phe$	55.95	>26.27	29.61	30.71	29.60	28.07	29.73	28.31
$U \rightarrow Pb+28Ne$	56.69	>26.27	30.48	31.23	30.27	28.98	28.63	29.53
$U \rightarrow Hg^{206}Hg^{-10}Mg$	70.73	>26.27	29.37	30.33	29.28	28.54	29.11	28.53
$U \rightarrow Hg^+ Mg$	72.27	>26.27	29.10	29.47	28.69	28.64	27.25	28.35
$U \rightarrow Hg^{+30}Mg$	69.46	-	32.81	33.98	32.72	30.91	30.69	31.71
$Np \rightarrow Bi + Ne$	61.90	_	20.61	21.95	20.10	21.18	24.32	19.53
$235 \times 207 \text{ m}; 282 \text{ f}$	62.16	_	22.05	23.24	21.46	22.64	23.98	22.10
237 Np $\rightarrow 207$ Tl+ 20 Mg	77.10	-	22.88	23.92	22.48	24.20	24.64	24.15
$Np \rightarrow 10^{10} \text{ T} + ^{10} \text{Mg}$	74.79	>27.57	27.17	28.63	26.82	27.13	27.22	27.79
$^{237}Pu \rightarrow ^{209}Pb + ^{29}Mg$	77.73	-	23.29	24.66	23.16	24.14	25.38	24.32
$Pu \rightarrow Pb + Mg$	77.45	_	24.43	25.73	24.18	24.95	25.68	25.01
$Pu \rightarrow Hg^{+}S_1$	91.46	-	25.29	26.48	25.39	25.43	26.47	26.35
$Pu \rightarrow Pb + Mg$	75.08	-	28.01	29.89	27.95	27.29	28.34	28.21
$Pu \rightarrow Hg + Si$	90.87	-	27.29	28.50	27.29	26.85	27.05	28.05
$Am \rightarrow B_1 + Mg$	79.85	-	22.15	23.35	22.02	23.02	24.52	22.54
237 Am $\rightarrow 207$ Tl+ 32 Si 241 $\rightarrow 207$ $\rightarrow 34$ $\rightarrow 34$	94.50	-	23.38	24.38	23.66	24.14	25.03	25.14
$Am \rightarrow C^{0}Tl + Si$	93.96	>24.41	25.26	26.26	25.44	25.51	25.54	26.77
$Cm \rightarrow Pb + Si$	97.55	-	21.64	21.09	22.02	22.87	22.24	22.24
$Cm \rightarrow {}^{209}Pb + {}^{2}Si$	95.39	-	23.66	25.02	24.26	24.07	25.63	25.25
$Cm \rightarrow Pb+Si$	94.79	-	25.61	27.00	26.12	25.47	26.19	26.92
$Cm \rightarrow Pb+Si$	93.17	-	27.24	27.89	27.92	26.43	26.34	27.12

 Table 3.
 Predicted half-lives for possible cluster radioactive nuclei.

B(A,Z) extracted from NUBASE2020 [52]. The relative sizes of the uncertainties of both the Q_c values and corresponding theoretical cluster radioactivity half-lives are also provided in Table 1. The good agreement between the experimental half-lives and the calculated values using our new G-N law presented in Table 1, Table 2, and Fig. 9 confirms that the new G-N law is reliable for cluster radioactivity half-life calculations. Therefore, we extend our new G-N law to predict the half-lives of 51 possible cluster radioactivity candidates, and the calculated results are displayed in Table 3. In this table, the labels of the first to eighth columns are similar to those of Table 1, except for the ninth column, which represents the predicted results of our new G-N law, denoted as $\lg T_{1/2}^{Cal}$. From Table 3, it is clear that the predicted results of our new G-N law are essentially consistent with the predicted values obtained using the compared formulas. In this table, we easily notice that in the case of $^{233}U \rightarrow$ ²⁰⁵Hg+²⁸Mg, the theoretical predicted half-lives calculated using all the analytical formulas are lower than the experimental bound by one to three orders of magnitude. A possible physical reason behind this result may be the effect of the proton subshell Z=92. Previous studies have discovered that the preformation probability near proton and neutron subshells strangely increase owing to the enhanced proton-neutron interactions, which results in a shorter half-life [57, 58]. This further confirms that

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cluster radioactivity is closely related to the nuclear structure. We hope that this study can be an effective tool toward theoretical investigation of cluster radioactivity and further provide a theoretical guide for future experiments.

IV. SUMMARY

In summary, based on Balasubramaniam's formula and considering the effect of the parent nucleus mass, blocking effect of unpaired nucleons, and reduced mass on cluster radioactivity half-lives, we propose a new G-N law for cluster radioactivity. The calculated cluster radioactivity half-lives obtained using our new G-N law are found to be in better agreement with experimental data with a least rms deviation of 0.606 compared with the results obtained using UNIV, the scaling law, EVS, Balasubramaniam's formula, and Ni's formula. The corresponding rms deviations are 0.829, 1.073, 0.744, 0.921, and 0.874, respectively. Furthermore, we extend our new G-N law to predict the cluster radioactivity half-lives of 51 nuclei whose decay energy are energetically allowed or observed but not yet quantified, and the decay energy Q_c is calculated from the latest 2020 atomic mass table. The predictions of our new G-N law are essentially consistent with those obtained using UNIV, the scaling law, EVS, Balasubramaniam's formula, and Ni's formula. The results of this study may be useful for probing the nuclear structure in future experiments.

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